

STARK BROADENING OF Se IV, Sn IV, Sb IV AND Te IV SPECTRAL LINES

M. S. Dimitrijević¹, Z. Simić¹, R. Stamm², J. Rosato²,
N. Milovanović¹ and C. Yubero³

¹*Astronomical Observatory, Volgina 7, 11060 Belgrade 38, Serbia*

²*Aix-Marseille Université, CNRS PIIM UMR 7345, 13397 Marseille Cedex 20, France*

³*Grupo de Física de Plasmas: Diagnósis, Modelos y Aplicaciones (FQM-136),
Universidad de Córdoba, Edificio A. Einstein (C-2), Campus de Rabanales,
14071 Córdoba, Spain*

*E-mail: mdimitrijevic@aob.rs, zsimic@aob.rs, roland.stamm@univ-amu.fr,
joel.rosato@univ-amu.fr, nmilovanovic@aob.rs, f62yusec@uco.es*

Stark broadening parameters, line width and shift, are needed for investigations, analysis and modelling of astrophysical, laboratory, laser produced and technological plasmas. Especially in astrophysics, due to constantly increasing resolution of satellite borne spectrographs, and large terrestrial telescopes, data on trace elements, insignificant before, now have increasing importance.

Using the modified semiempirical method (Dimitrijević and Konjević, 1980), here have been calculated Stark widths for 1 Se IV, 5 Sn IV, 2 Sb IV and 1 Te IV transitions. Since Sn, Sb, and Te are successive elements in the periodic system, and Se and Te homologous, the obtained results have been used to discuss systematic trends.

Obtained results will be implemented in the STARK-B database (<http://stark-b.obspm.fr>) which is also a part of Virtual atomic and molecular data center (VAMDC - <http://www.vamdc.org/>).

References

Dimitrijević, M. S., Konjević, N.: 1980, Stark widths of doubly- and triply-ionized atom lines, *Journal of Quantitative Spectroscopy and Radiative Transfer*, **24**, 451.